

# Cosmology Simulations with Enzo

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Enzo Workshop  
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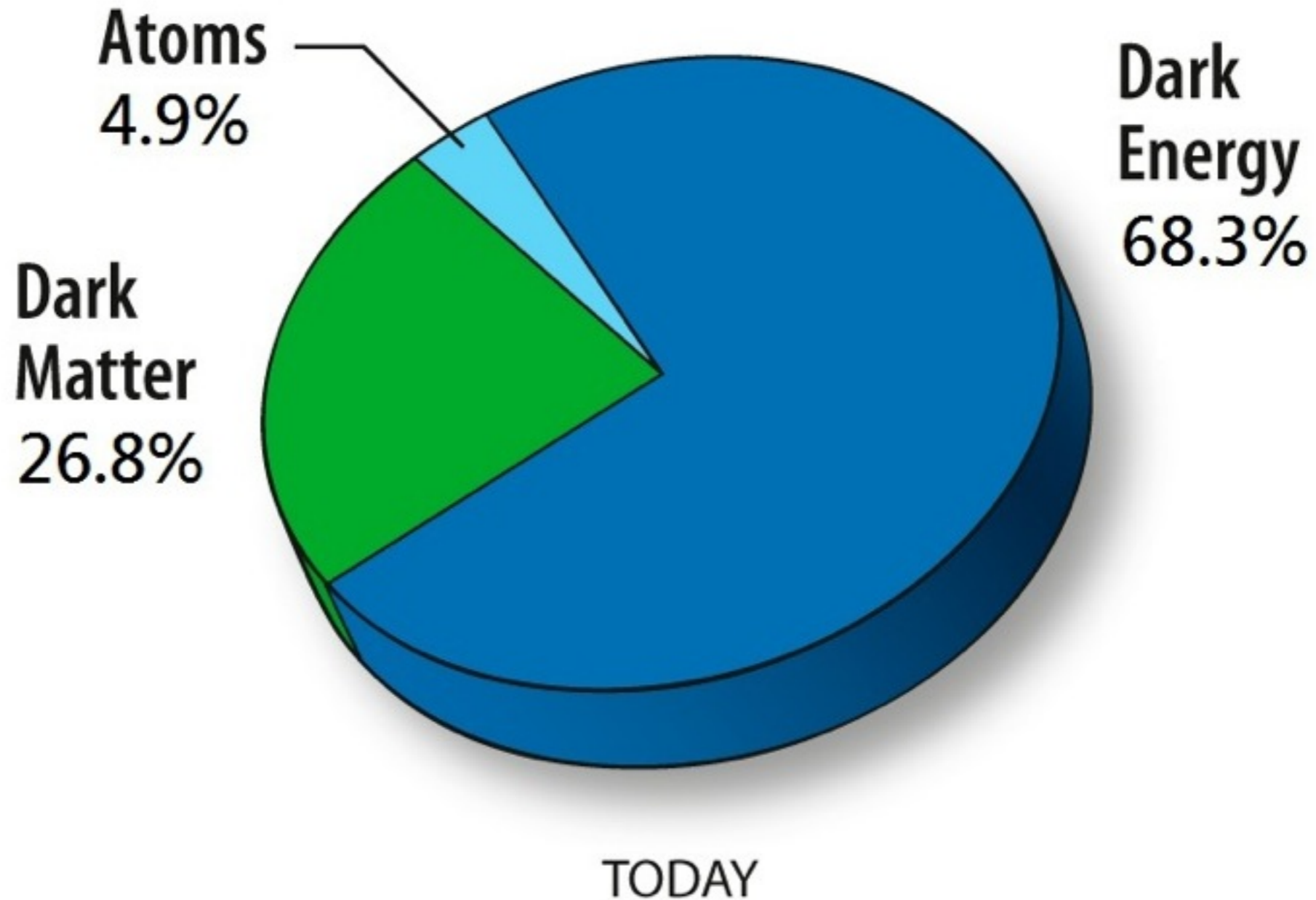
# Outline

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- Introduction to unigrid cosmology simulations
- Introduction to nested grid cosmology simulations
- Using different non-equilibrium chemistry models
- Including radiative cooling & star formation/feedback
- Including radiative transfer

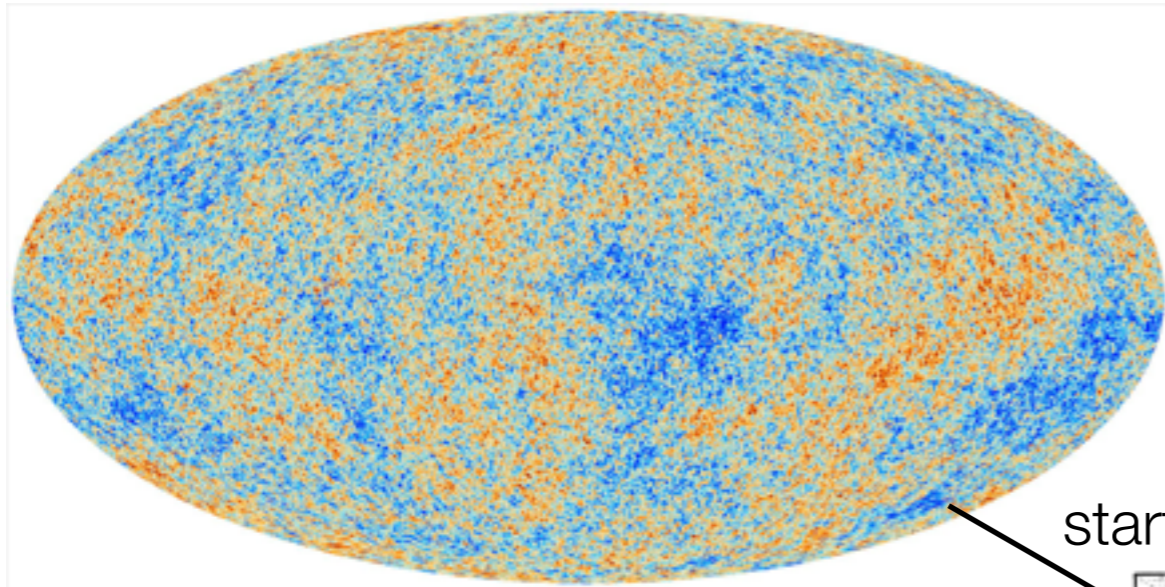
# A universe in a box

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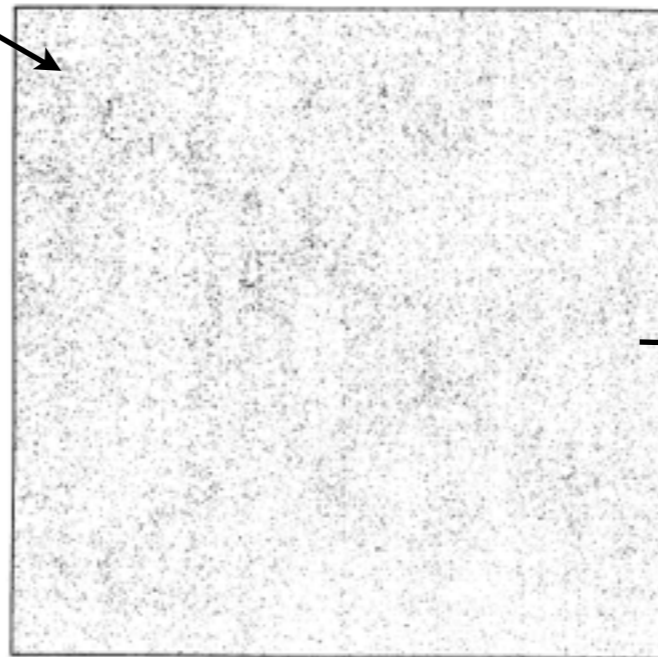
# A universe in a box

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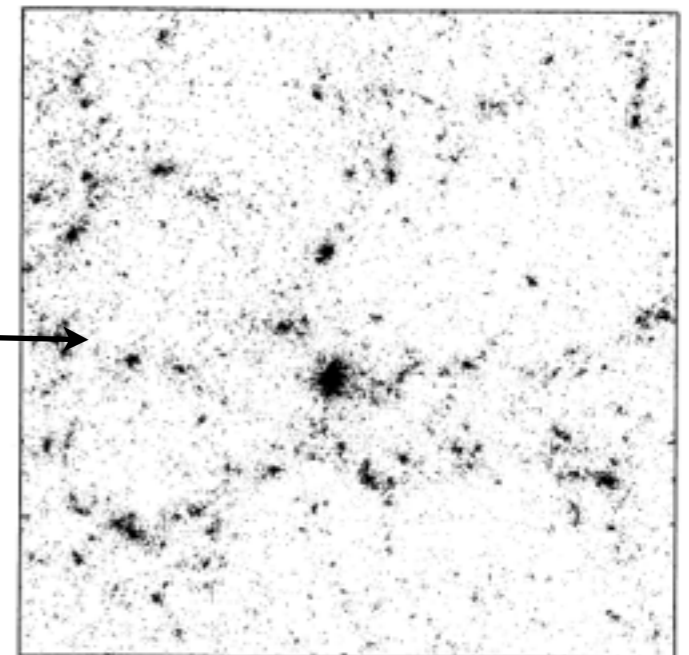


$z \sim 1080$

Linearly evolve to a starting redshift ( $z \sim 50-200$ )



White et al. (1985)



Use simulations to evolve in the non-linear regime

# Unigrid Initial Conditions

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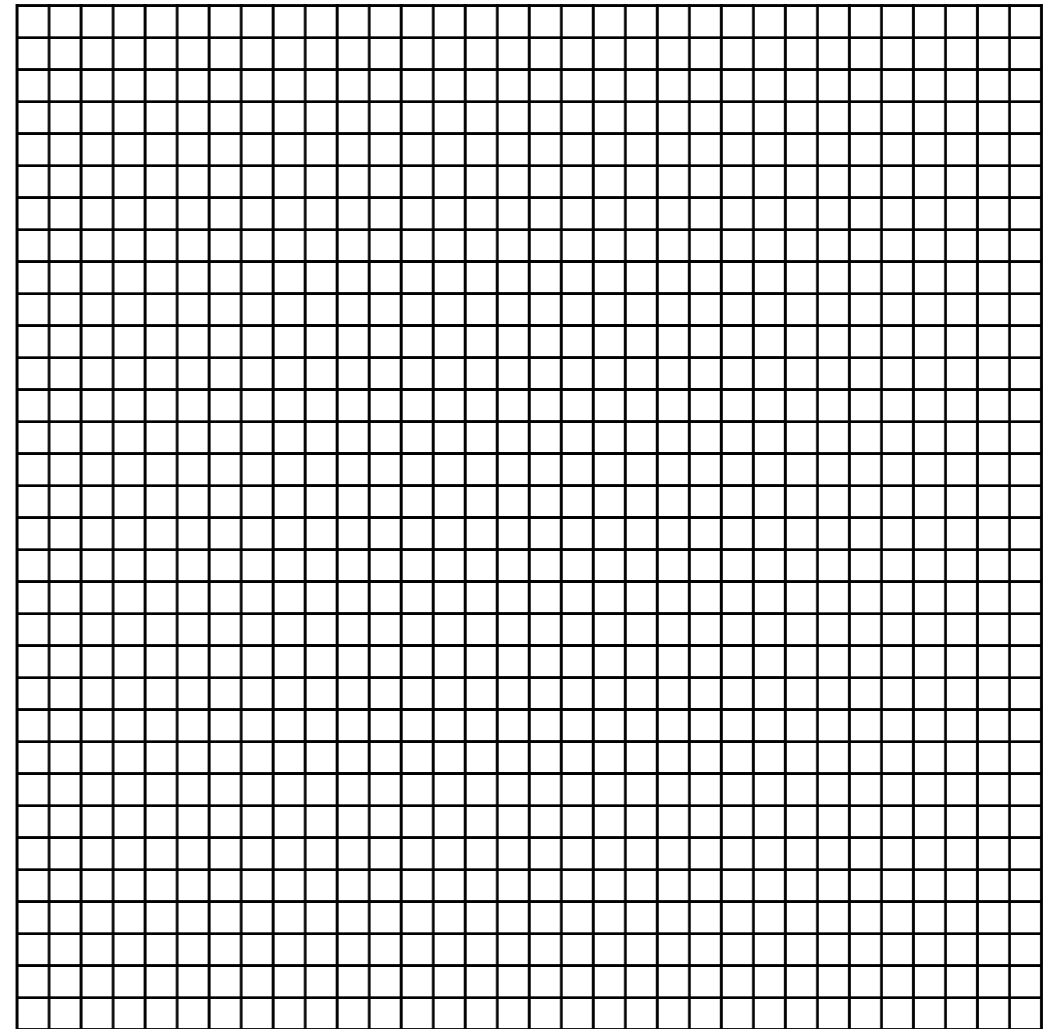
- Two programs are provided with Enzo to generate cosmology initial conditions: **inits** (src/inits) or **mpgrafic** (src/mpgrafic)
- However, we will be using the program **MUSIC** (Hahn & Abel 2011) that is more robust and efficient.
  - <http://www.phys.ethz.ch/~hahn/MUSIC/index.html>
  - <https://bitbucket.org/ohahn/music>
- 10 initial condition files in HDF5 file format
  - Baryon Density:  $1 \times N \times N \times N$
  - (x,y,z) Baryon Velocity:  $1 \times N \times N \times N$
  - (x,y,z) Particle Velocity:  $1 \times N \times N \times N$
  - (x,y,z) Particle Displacement:  $1 \times N \times N \times N$



# Unigrid Initial Conditions with MUSIC

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- MUSIC: MUlti-Scale Initial Conditions
- Designed specifically for nested grid “zoom-in” simulations.
- Improves errors by two orders of magnitude compared to previous codes.
- Zel’dovich approximation and 2LPT
- Multi-code support (e.g. GADGET, ART, Topsy, RAMSES, Enzo)
- Publicly available, OpenMP parallel



# Unigrid Initial Conditions with MUSIC

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- **Requirements:**
  - **GNU Scientific Library (GSL)**
  - **FFTW 2 or 3**
  - **HDF5**

# Installing MUSIC

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- Download the source code (on conival)
  - `hg clone https://bitbucket.org/ohahn/music`
  - `cp ~guest01/music/Makefile music/`
  - `cd music`
  - `make`
  - `export LD_LIBRARY_PATH=${LD_LIBRARY_PATH}:/home/guest01/local/lib`



# Grab MUSIC and Enzo parameter files

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- On conival,
  - `cd ~`
  - `tar xf ~guest01/sapporo_cosmo.tar`
- On your laptop
  - `wget http://www.physics.gatech.edu/~jw254/scpics/sapporo\_cosmo.tgz`
  - `tar xf sapporo_cosmo.tgz`

# Initial Conditions Parameters

---

- Cosmology Parameters: box size, Hubble constant, mass-energy fractions ( $\Omega_\Lambda$ ,  $\Omega_m$ ,  $\Omega_b$ ), initial redshift
- Power spectrum parameters: power spectrum type,  $\sigma_8$ , random seed

```
[setup]
boxlength      = 5
zstart         = 50
```

units of comoving Mpc/h

level = 5  $\rightarrow$   $32^3$  resolution

```
[cosmology]
Omega_m        = 0.3175
Omega_L        = 0.6825
Omega_b        = 0.049
H0             = 67.11
sigma_8        = 0.8344
nspec          = 0.9624
transfer       = eisenstein

[random]
seed[5]        = 201310160
```

# Initial Conditions Parameters – Grid parameters

---

```
[setup]
boxlength      = 5
zstart        = 50
levelmin      = 5
levelmin_TF   = 5
levelmax      = 5
baryons       = yes
use_2LPT      = yes
use_LLA       = yes
periodic_TF   = yes
```

Initial redshift

Base resolution =  $2^n$  cells across

Finest resolution of zoom-in region

Generate baryon ICs

2LPT = yes; Zel'dovich = no

```
MUSIC% ~/codes/music/MUSIC sapporo_cosmo.conf
```



```
this is music! versio
```

- Opening log file 'sapporo\_cosmo.conf\_log.txt'
- Selecting transfer function plug-in 'eisenste
- starting at  $a=0.0196078$
- Selecting region generator plug-in 'box'...
- WARNING: The selected transfer function does not support distinct amplitudes for baryon and DM. Perturbation amplitudes will be identical.

```
- Grid structure:  
Level 5 : offset = ( 0, 0, 0 )  
           size = ( 32, 32, 32 )
```

```
- Finest level :  
   extent = 5 x 5 x 5 h-3 Mpc**3  
   mtotgrid = 1.10141e+13 h-1 M_sun  
   particle mass = 2.84249e+08 h-1 M_sun  
   baryon mass/cell = 5.1874e+07 h-1 M_sun  
   dx = 0.15625 h-1 Mpc
```

```
- Selecting output plug-in 'enzo'...
```

```
GENERATING WHITE NOISE
```

- Density calculation took 0.170149s with 4 threads.
- Top grid mean density is off by  $-4.1901e-17$ , correcting...
- Global density extrema:
  - minimum:  $\delta=-0.583318$  at (0.765625,0.953125,0.515625)  
shifting back at (0.765625,0.953125,0.515625)
  - maximum:  $\delta=0.664508$  at (0.359375,0.796875,0.921875)  
shifting back at (0.359375,0.796875,0.921875)
- Invoking unigrid FFT Poisson solver...
- Invoking unigrid FFT Poisson solver...

```
COMPUTING BARYON DENSITY
```

- Using real-space transfer function kernel.
  - Computing transfer function kernel...
  - Deconvolving density kernel...
  - Performing noise convolution on level 5 ...
- Fetching kernel for level 5
  - Performing density convolution... (32, 32, 32)
- Density calculation took 0.1516s with 4 threads.
- Top grid mean density is off by  $-4.19006e-17$ , correcting...
- Invoking unigrid FFT Poisson solver...
- Invoking unigrid FFT Poisson solver...
- Top grid mean density is off by 0.0175262, correcting...
- Wrote output file 'sapporo\_cosmo' using plugin 'enzo'...
- Done!

# Set up *Enzo* parameter file

---

- MUSIC creates a basic but non-functional parameter file, `parameters.txt`.
- The user must add the relevant AMR, hydro, chemistry, etc. parameters.
- See `run/Cosmology/Hydro/AMRCosmology/AMRCosmology.enzo` for an example.
- For this workshop, we have provided a set of example parameter files for each lab exercise.

# Run the simulation

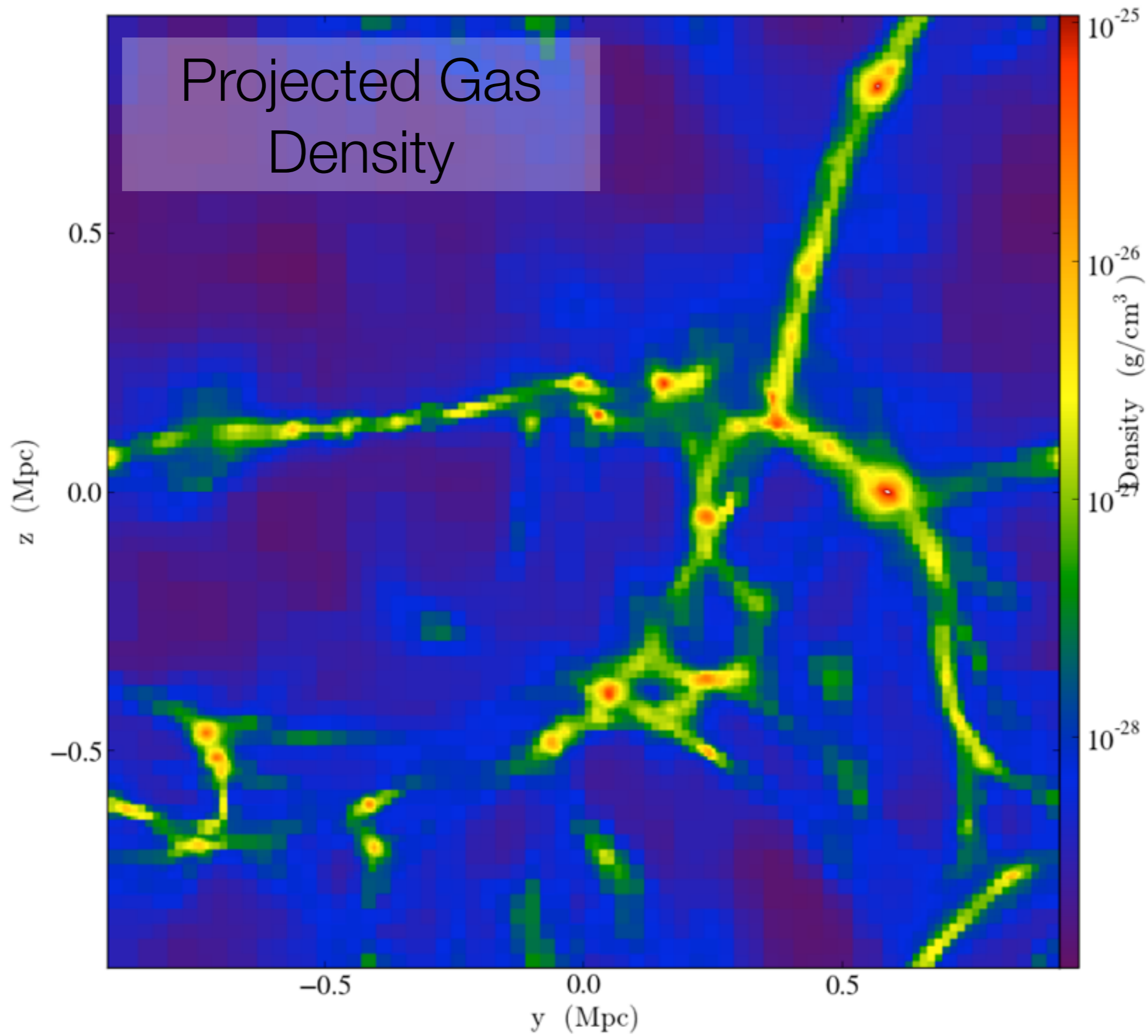
---

- Move parameter file and initial conditions into the same directory (see `Adiabatic` directory)
- Run enzo!

```
sapporo_cosmo% ./enzo.exe adiab.enzo
```

- Add the “-d” option for debugging information
- Add the “-r” option for restarting from a snapshot.

```
sapporo_cosmo% ./enzo.exe -d -r DD0010/DD0010
```

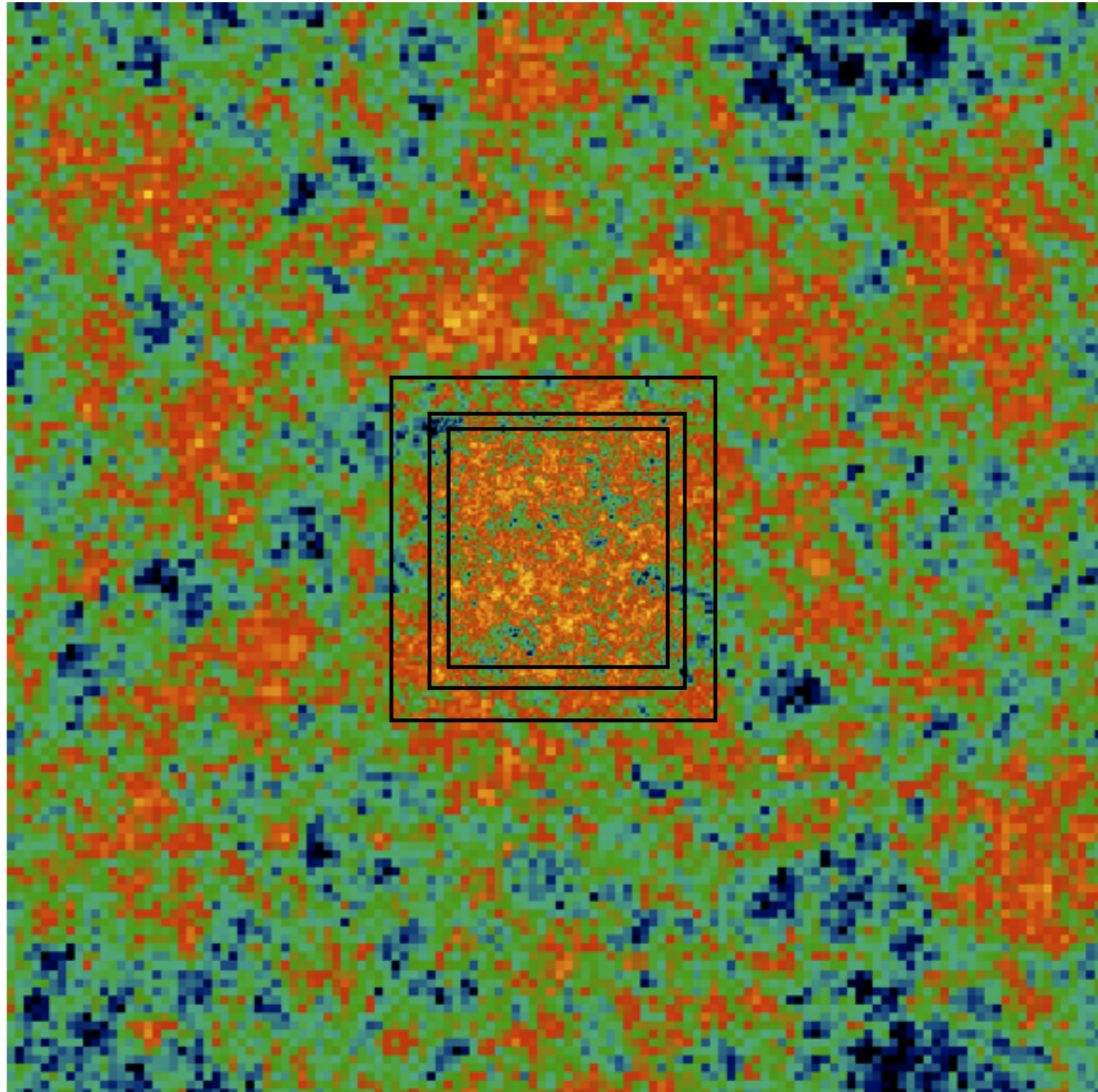


# Nested Grid Cosmology Simulations

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aka *Zoom-in* calculations

1. *N*-Body Simulation
2. Locate region of interest, e.g. most massive halo
3. Generate nested grid initial conditions
4. Run simulation!





# N-body Simulation

---

- Usually run at the same resolution as the production level-0 resolution
- Create initial conditions with **MUSIC** and use\_baryons=no
- Key parameters

```
CosmologySimulationOmegaBaryonNow      = 0.0
CosmologySimulationOmegaCDMNow         = 0.3175
CosmologySimulationParticleVelocity1Name = ParticleVelocities_x
CosmologySimulationParticleVelocity2Name = ParticleVelocities_y
CosmologySimulationParticleVelocity3Name = ParticleVelocities_z
CosmologySimulationParticleDisplacement1Name = ParticleDisplacements_x
CosmologySimulationParticleDisplacement2Name = ParticleDisplacements_y
CosmologySimulationParticleDisplacement3Name = ParticleDisplacements_z
CosmologySimulationNumberOfInitialGrids = 1
CosmologySimulationCalculatePositions = 1
```

```
RefineBy          = 2 // refinement factor
#CellFlaggingMethod = 2 4 // use baryon and DM mass
CellFlaggingMethod = 4 // use DM mass for refinement
MinimumEfficiency = 0.4 // fraction efficiency
MinimumOverDensityForRefinement = 4. // times the initial dens
```

# Halo finding

---

- On the last output, we need to find the cosmological halos.
- Enzo has a built-in halo finder. Run it with

```
sapporo_cosmo% ./enzo.exe -F -d -r DD0020/DD0020
```

- Will put the halo list sorted by mass in the FOF/ directory.

```
# Time = 15.921139
# Redshift = 12.802003
# Number of halos = 83
#
# Column 1. Center of mass (x)
# Column 2. Center of mass (y)
# Column 3. Center of mass (z)
# Column 4. Halo number
# Column 5. Number of particles
# Column 6. Halo mass [solar masses]
# Column 7. Virial mass [solar masses]
# Column 8. Stellar mass [solar masses]
# Column 9. Virial radius (r200) [kpc]
# Column 10. Mean x-velocity [km/s]
# Column 11. Mean y-velocity [km/s]
# Column 12. Mean z-velocity [km/s]
# Column 13. Velocity dispersion [km/s]
# Column 14. Mean x-angular momentum [Mpc * km/s]
# Column 15. Mean y-angular momentum [Mpc * km/s]
# Column 16. Mean z-angular momentum [Mpc * km/s]
# Column 17. Spin parameter
#
# datavar lines are for partiview. Ignore them if you're not partiview.
#
datavar 0 halo_number
datavar 1 number_of_particles
datavar 2 halo_mass
datavar 3 virial_mass
datavar 4 stellar_mass
groups_00137.dat lines 1-29/122 4%
```

```

datavar 0 halo_number
datavar 1 number_of_particles
datavar 2 halo_mass
datavar 3 virial_mass
datavar 4 stellar_mass
datavar 5 virial_radius
datavar 6 x_velocity
datavar 7 y_velocity
datavar 8 z_velocity
datavar 9 velocity_dispersion
datavar 10 x_angular_momentum
datavar 11 y_angular_momentum
datavar 12 z_angular_momentum
datavar 13 spin

```

Position

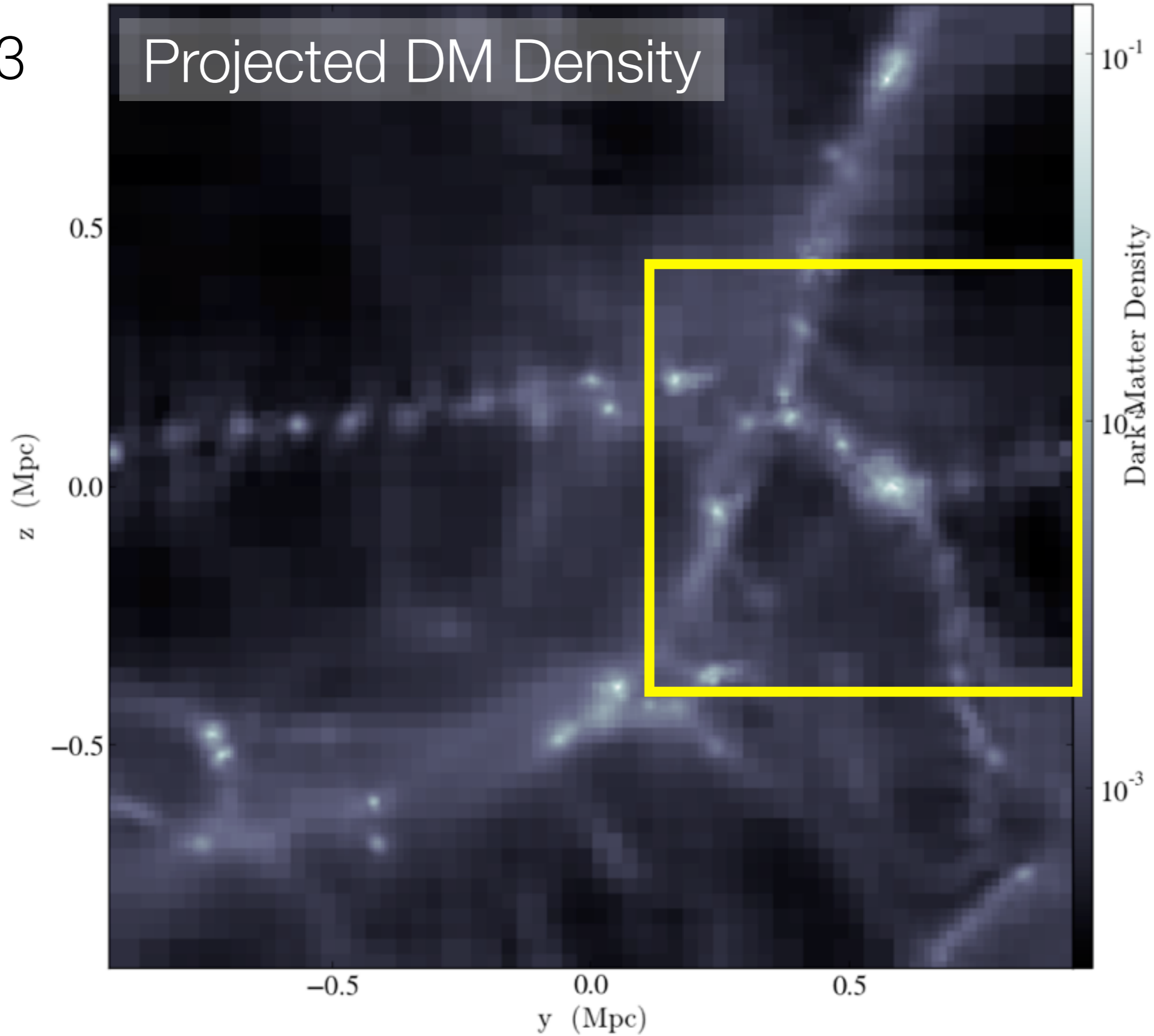
Mass

0.76364088	0.29754913	0.37248853	0	2060	571700.88
447925.16	0	0.11323795	0.84548277	-0.41380891	1.0349354
5.341588	2.408008e-05	3.4065459e-05	-6.1198812e-06	0.11691254	
0.46177065	0.61586523	0.6905852	1	1390	385758.56
372992.47	0	0.10630372	2.0898418	0.23426482	-1.659736
.6082692	5.5966311e-06	1.264129e-05	-3.7420041e-06	0.041144241	
0.84373128	0.67770523	0.27474043	2	415	115172.2
70213.453	0	0.060892154	-0.90891731	-0.64614326	2.0012767
.8843138	2.717258e-06	-1.8763769e-06	-7.576411e-06	0.078942753	
0.51242203	0.62618536	0.67746228	3	372	103238.98
95745.844	0	0.067207284	-0.48392525	0.49669054	-2.3942778
.9666159	7.2602943e-06	3.2799553e-06	-2.7749318e-06	0.060778547	
0.77068067	0.30860633	0.33854923	4	330	91583
81314.617	0	0.06338729	-1.6947181	-0.76077771	3.8264091

```
groups_00137.dat lines 25-44/122 11%
```

$z = 3$

# Projected DM Density



# Nested Grid Initial Conditions with MUSIC

---

- Find halo of interest.
- We want to re-sample the Lagrangian volume of this region with a higher resolution.
- We can use yt to find this volume. Sample script can be found at
- [https://bitbucket.org/MatthewTurk/yt.lagrangian\\_volume](https://bitbucket.org/MatthewTurk/yt.lagrangian_volume)
  - All particles have unique IDs, so we can find the particles in some region in the final output.
  - Then we find the bounding box of the same particles in the initial output.

# MUSIC: Nested Grid Parameters

```
[setup]
boxlength      = 5
zstart        = 50
levelmin      = 5
levelmin_TF   = 5
levelmax      = 7
padding       = 2
overlap       = 0
ref_center    = 0.3, 0.81, 0.5
ref_extent    = 0.25, 0.25, 0.25
align_top     = yes
baryons       = no
use_2LPT     = yes
use_LLA      = yes
periodic_TF   = yes
```

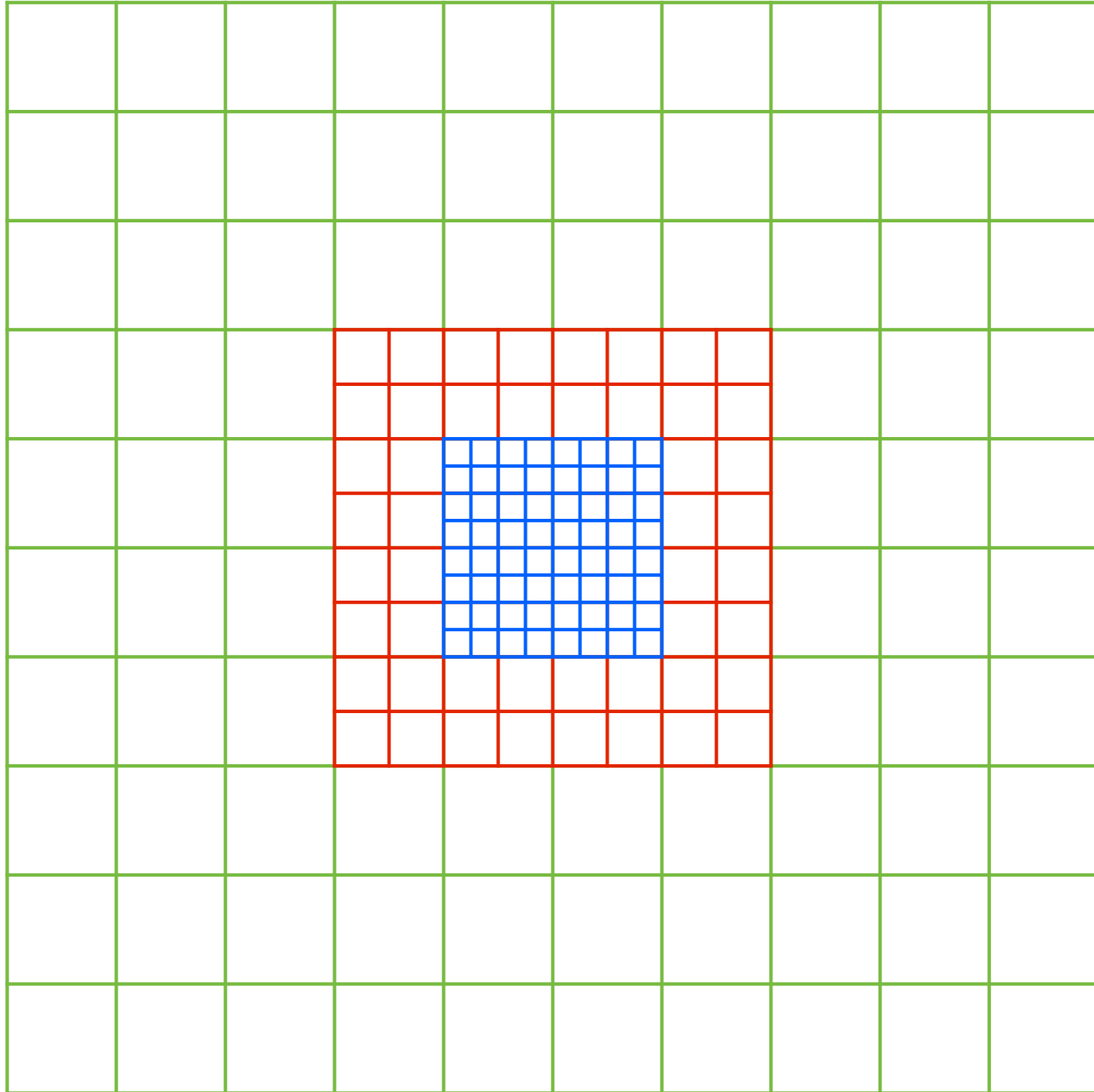
Resolution of the finest nested grid.  
Effective resolution =  $2^7 = 128$

Center of the Lagrangian volume  
Width of the Lagrangian volume

```
[random]
seed[5]       = 201310160
seed[6]       = 301310160
seed[7]       = 401310160
```

Optional: New random seeds for nested grids.

padding = 2





```
sapporo_cosmo% ~/codes/music/MUSIC sapporo_cosmo_nested.conf
```

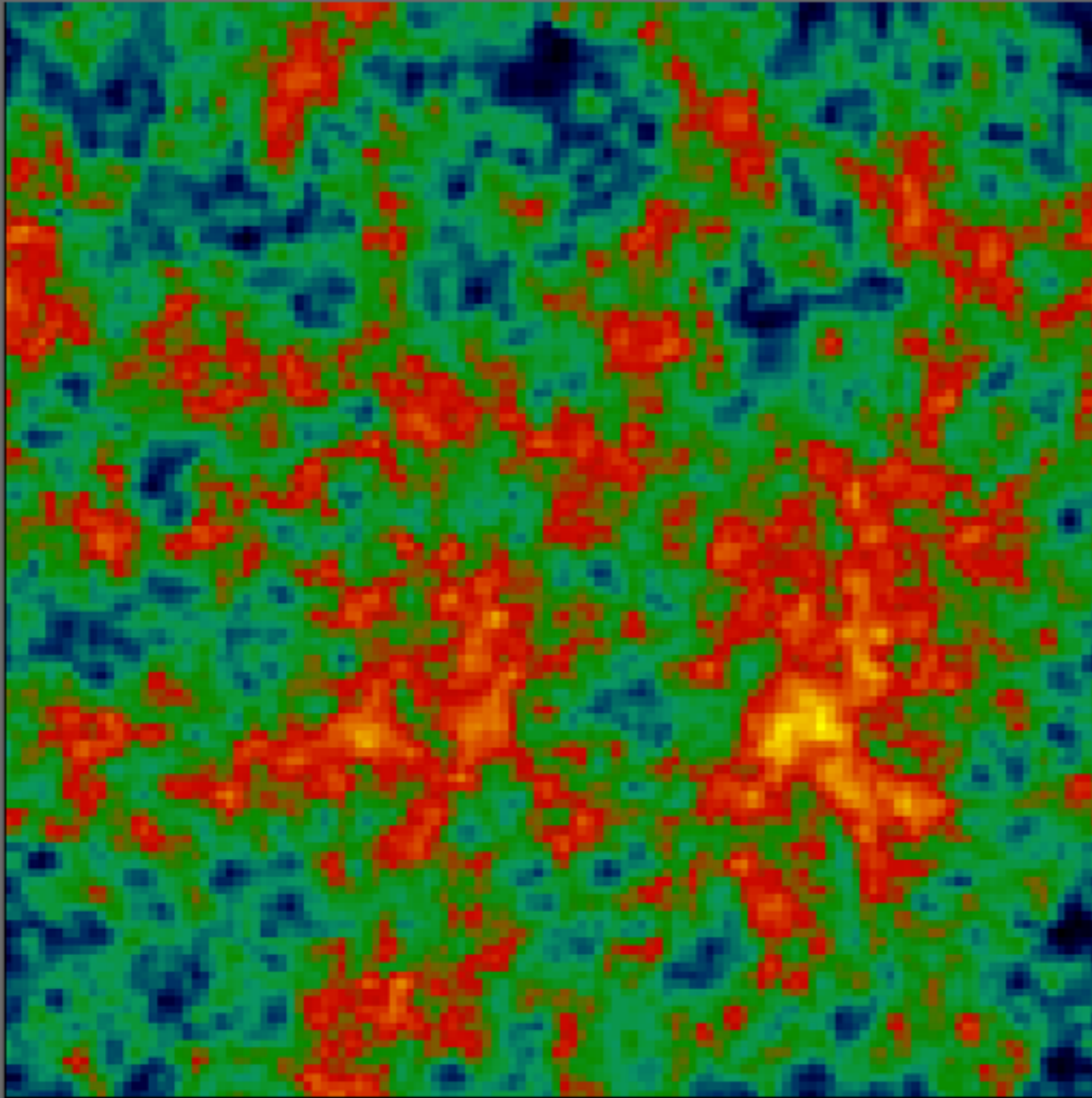


# with MUSIC

this is mus

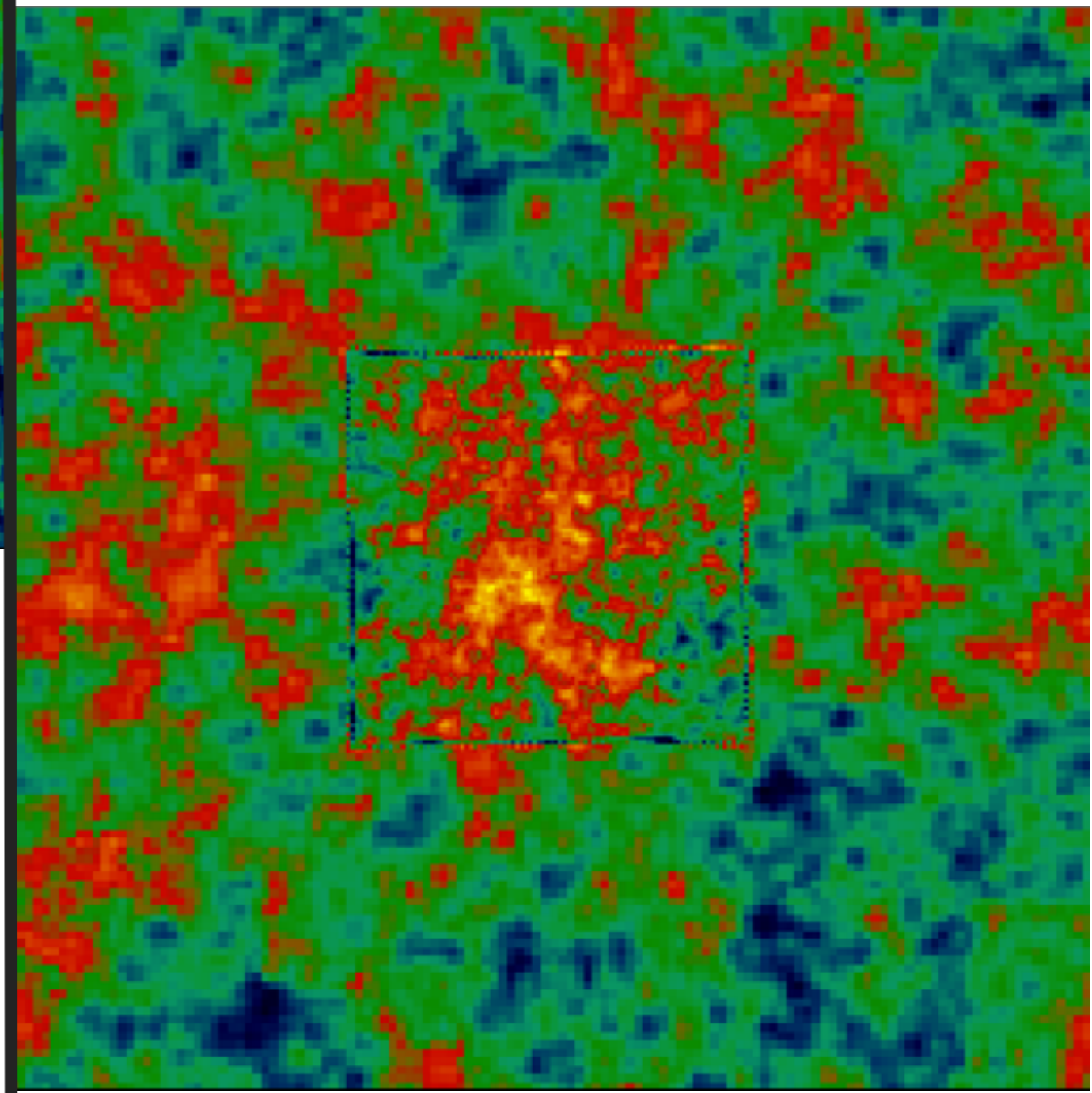
- Opening log file 'sapporo\_cosmo\_nest
  - Selecting transfer function plug-in
  - starting at  $a=0.0196078$
  - Selecting region generator plug-in '
  - refinement region is 'box', w/ bound  
left = [0.175000,0.685000,0.373  
right = [0.425000,0.935000,0.626
- 
- Domain will be shifted by (6, -9, 0)
- Grid structure:
    - Level 5 : offset = ( 0,
    - size = ( 32,
    - Level 6 : offset = ( 9,
    - size = ( 26,
    - Level 7 : offset = ( 2,
    - size = ( 40,
- 
- Finest level :
    - extent = 1562.5 x 1
    - mtotgrid = 3.36123e+1
    - particle mass = 5.25192e+0
    - dx = 39.0625 h-
- 
- Selecting output plug-in 'enzo'...

- Top grid mean density is off by  $-6.5921e-05$ , correcting...
  - Global density extrema:
    - minimum: delta= $-0.730190$  at (0.449219,0.378906,0.425781) (level=7)
    - shifted back at (0.261719,0.660156,0.425781)
    - maximum: delta= $0.986598$  at (0.582031,0.621094,0.488281) (level=7)
    - shifted back at (0.394531,0.902344,0.488281)
  - Density extrema on finest level:
    - minimum: delta= $-0.730190$  at (0.449219,0.378906,0.425781)
    - shifted back at (0.261719,0.660156,0.425781)
    - maximum: delta= $0.986598$  at (0.582031,0.621094,0.488281)
    - shifted back at (0.394531,0.902344,0.488281)
- 
- Invoking multi-grid Poisson solver...
  - Converged in 5 steps to  $3.89154e-06$
  - Poisson solver took 0.742935s with 4 threads.
  - Computing 2LPT term....
  - Solving 2LPT Poisson equation
- 
- Invoking multi-grid Poisson solver...
  - Converged in 5 steps to  $2.05267e-06$
  - Poisson solver took 0.775615s with 4 threads.
    - Performing hybrid Poisson step... (80, 80, 80)
    - velocity component 0 : sigma = 3.0004
    - velocity component 1 : sigma = 1.79269
    - velocity component 2 : sigma = 1.84676
- 
- COMPUTING DARK MATTER DISPLACEMENTS
- 
- Performing hybrid Poisson step... (80, 80, 80)
- 
- Wrote output file 'IC.2'  
          using plugin 'enzo'...
  - Done!



Unigrid  
Smoothed DM Density Slice

1 Nested Grid



# Setting up the Nested Grid Parameter File

```
CosmologySimulationNumberOfInitialGrids = 3
CosmologySimulationGridDimension[1] = 26 26 26
CosmologySimulationGridLeftEdge[1] = 0.28125 0.34375 0.28125
CosmologySimulationGridRightEdge[1] = 0.6875 0.75 0.6875
CosmologySimulationGridLevel[1] = 1
CosmologySimulationGridDimension[2] = 40 40 40
CosmologySimulationGridLeftEdge[2] = 0.3125 0.375 0.3125
CosmologySimulationGridRightEdge[2] = 0.625 0.6875 0.625
CosmologySimulationGridLevel[2] = 2
#
# region allowed for further refinement
#
RefineRegionLeftEdge = 0.3125 0.375 0.3125
RefineRegionRightEdge = 0.625 0.625 0.625
"
```

Nested Grid Dimensions  
and Boundaries

Refine region restricts AMR grids within these coordinates.  
Should equal or be smaller than the finest nested grid

# Overdensity criteria

---

- **Very important:** Must divide the baryon and DM mass refinement criteria by a factor of  $8^n$ , where  $n :=$  maximum initial level
- e.g. Refine on an overdensity of 3 with 2 initial grids  $\rightarrow 3.0/64 = 0.046875$
- If not done, the hierarchy will remain under-resolved, compared with the desired results.

```
#
# set grid refinement parameters
#
StaticHierarchy          = 0      // dynamic hierarchy
MaximumRefinementLevel  = 8
MaximumGravityRefinementLevel = 8
MaximumParticleRefinementLevel = 8
RefineBy                 = 2      // refinement factor
CellFlaggingMethod      = 2 4 6   // use baryon mass for refinement
#CellFlaggingMethod     = 4       // use DM mass for refinement
MinimumEfficiency       = 0.3     // fraction efficiency
MinimumOverDensityForRefinement = 0.5 0.5 // times the initial density refers to top grid: divide by 8 for each additional level
MinimumMassForRefinementLevelExponent = -0.2 0.0
```

# Run enzo! (see directory: DarkMatter-Nested)

---

- Running on a single core, outputting standard output to display.

```
sapporo_cosmo% ./enzo.exe -d nbody_nested.enzo
```

- Running on 4 cores, piping standard output to the file, estd.out.

```
sapporo_cosmo% mpirun -n 4 ./enzo.exe -d nbody_nested.enzo >& estd.out
```

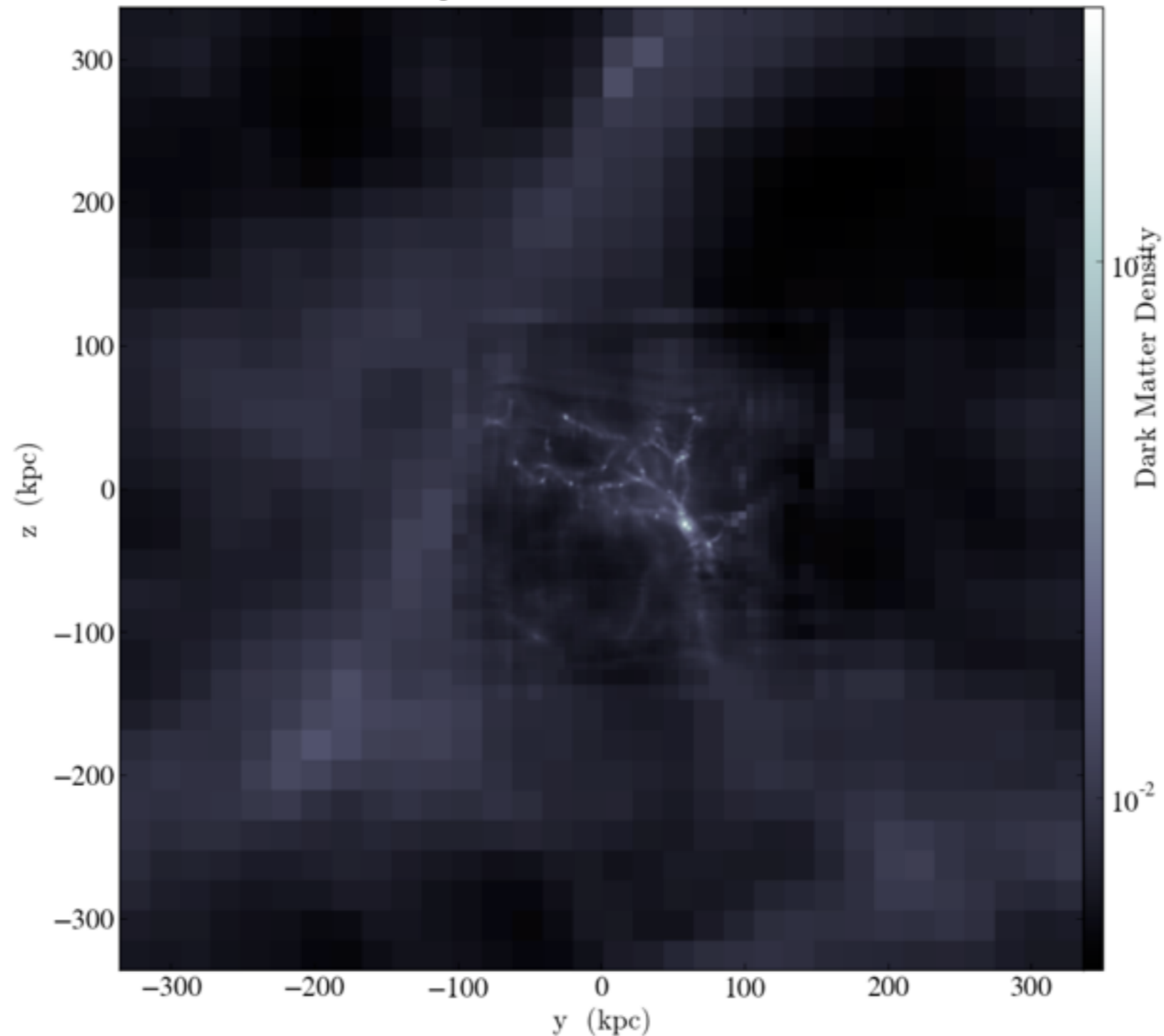
- This simulation takes significantly longer than the simulation with no nested grids. On my laptop, it took 20 minutes on 2 cores.

# Differences between simulations

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- Same spatial resolution but higher mass resolution!

$z = 10$

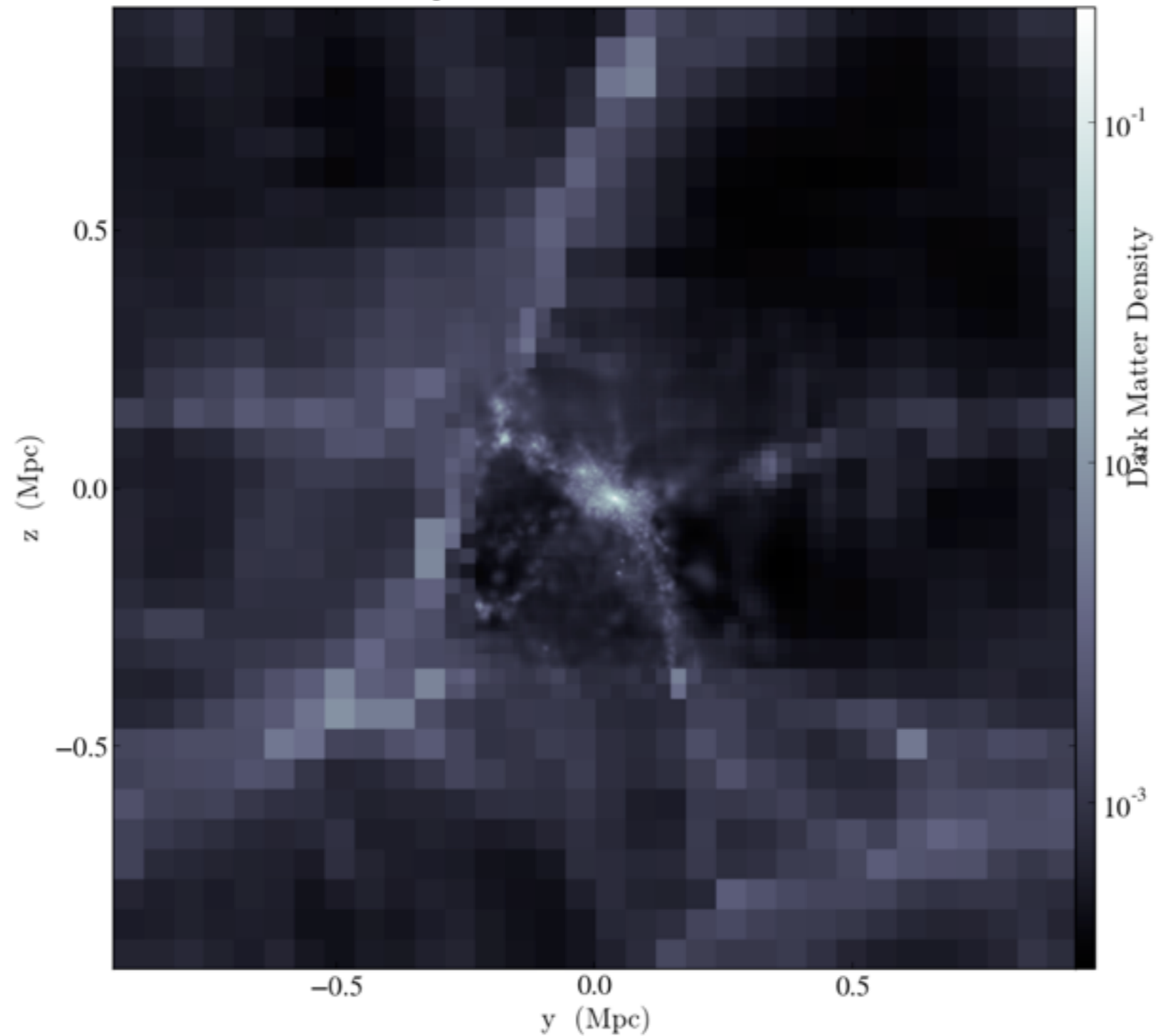


# Differences between simulations

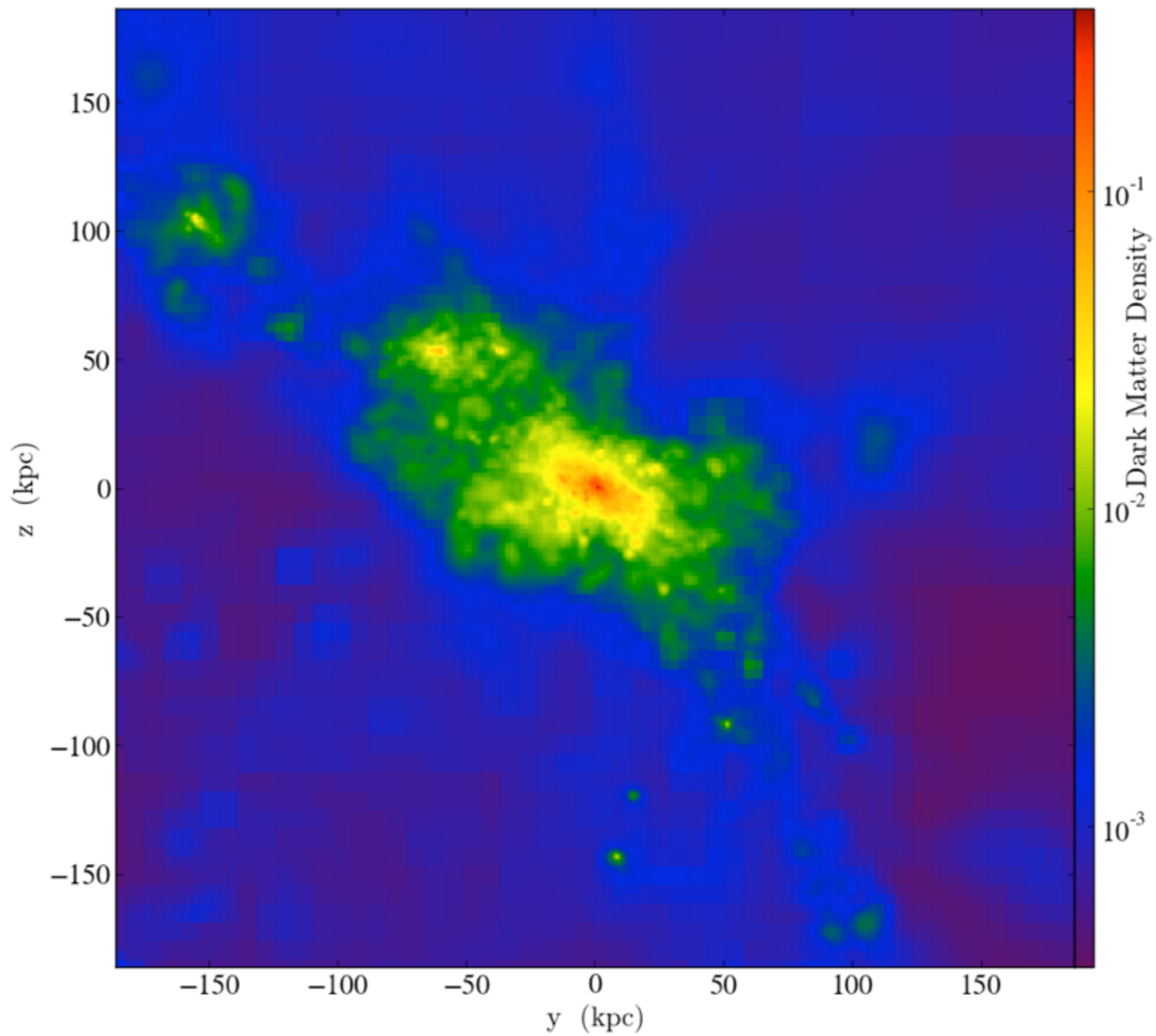
---

- Same spatial resolution but higher mass resolution!

$z = 3$



$z = 3$





# Different chemistry and cooling models

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- **MultiSpecies == 0:** Equation of state with an adiabatic index  $\gamma$ 
  - **RadiativeCooling == 1:** Cooling from a tabulated cooling curve (copy input/cool\_rates.in to working directory)

## Non-equilibrium chemistry

- **MultiSpecies == 1:** 6-species (H, H<sup>+</sup>, He, He<sup>+</sup>, He<sup>++</sup>, e<sup>-</sup>)
- **MultiSpecies == 2:** 9-species (H, H<sup>+</sup>, He, He<sup>+</sup>, He<sup>++</sup>, e<sup>-</sup>, H<sup>-</sup>, H<sub>2</sub>, H<sub>2</sub><sup>+</sup>); i.e. **+molecular hydrogen cooling**
- **MultiSpecies == 3:** 12-species (H, H<sup>+</sup>, He, He<sup>+</sup>, He<sup>++</sup>, e<sup>-</sup>, H<sup>-</sup>, H<sub>2</sub>, H<sub>2</sub><sup>+</sup>, D, D<sup>+</sup>, HD); **+HD cooling**

# Different chemistry and cooling models

---

- **MetalCooling == 1:** Tabulated cooling table from Glover & Jappsen (2007) & Sutherland & Dopita (1993) – input/metal\_cool.dat.
  - $T < 10^4$  K: fine-structure metal-line cooling
  - $T > 10^4$  K: collisional ionization equilibrium
- **MetalCooling == 2:** Explicitly calculated metal cooling rates (Cen)
- **MetalCooling == 3:** Tabulated cooling table from CLOUDY (Smith et al. 2008)

## 3 more parameters for **large** nested grid runs

---

- i.e.  $>256^3$  nested grids
- `ParallelRootGridIO = 1`: Root grids are tiled before I/O, and each processor reads/writes their own tile
- `PartitionNestedGrids = 1`: Same thing as above but for nested grids
- `UnigrdTranspose = 2`: More parallel-efficient transpose for the FFT done for the gravity solver.

# Summary

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## You should now know **how to**:

- **Create** cosmological unigrid and nested grid initial conditions with MUSIC. For more, see
  - <http://www.phys.ethz.ch/~hahn/MUSIC/index.html>
  - <https://bitbucket.org/ohahn/music>
  - [https://bitbucket.org/ohahn/music/downloads/MUSIC\\_Users\\_Guide.pdf](https://bitbucket.org/ohahn/music/downloads/MUSIC_Users_Guide.pdf)
- **Run** cosmological regular and nested grid runs ( $N$ -Body and  $N$ -Body + gas)
  - <http://enzo.readthedocs.org/en/latest/tutorials/RunCosmologySimulation.html>
  - However, these instructions are for the *inits* initial conditions generator that is included with Enzo.